

RADAER AEROSOL-RADIATION INTERACTIONS



Nicolas Bellouin

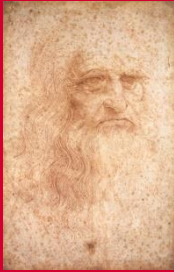
UKCA Training Workshop, Cambridge, 11 January 2017

LECTURE SUMMARY

- **Why care about aerosol-radiation interactions?**
- **Theory of aerosol-radiation interactions**
 - **Mie scattering**
- **Description of RADAER**
 - **Methods**
 - **Diagnostics**
 - **Double-call simulations**

AEROSOL-RADIATION INTERACTIONS

VERY SHORT HISTORY



~1490

Leonardo da Vinci notes that the side of a dust and smoke plume facing the Sun is far brighter than the other side.



1908

Gustav Mie publishes in *Annalen der Physik* the solution to Maxwell equations for a homogeneous dielectric sphere.



1971

Stephen Schneider and S.I. Rasool estimate in *Science* that anthropogenic aerosol cooling will dominate CO₂ warming.

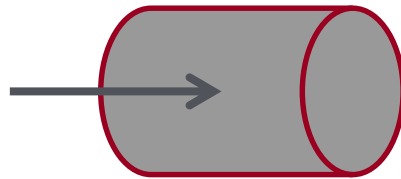


1980s

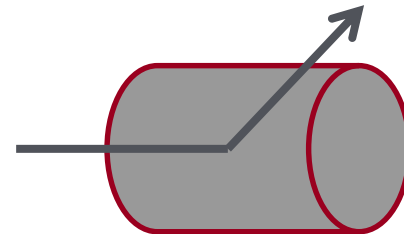
Atsumu Ohmura reports sizeable decreases in solar radiation reaching the surface, a phenomenon later coined “global dimming” and linked to aerosol-radiation interactions.

AEROSOL-RADIATION INTERACTIONS

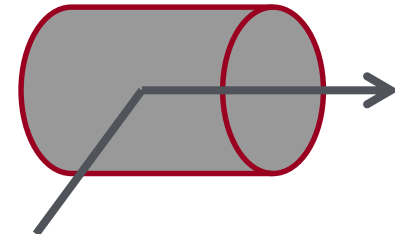
- Absorption



- Scattering out of viewing direction



- Scattering into viewing direction



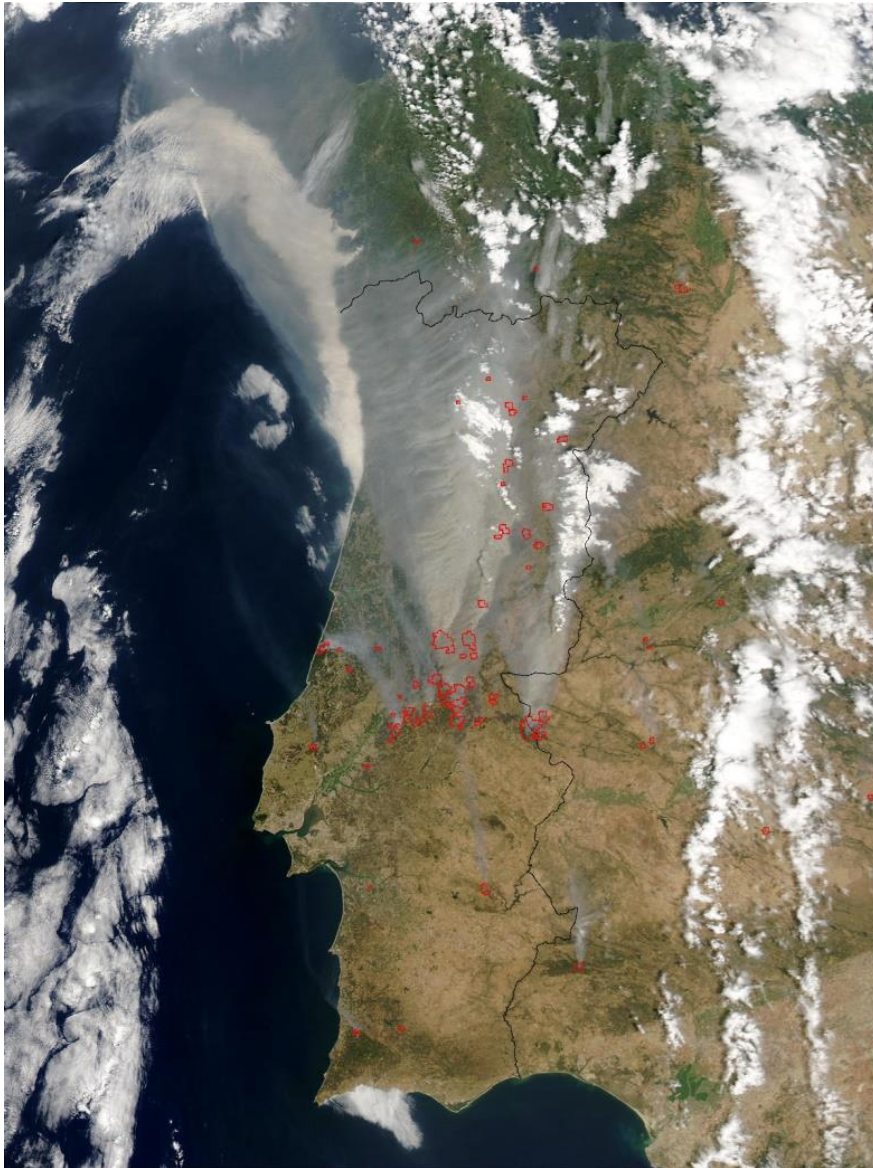
- Emission



WHY WE CARE ABOUT AEROSOL-RADIATION INTERACTIONS

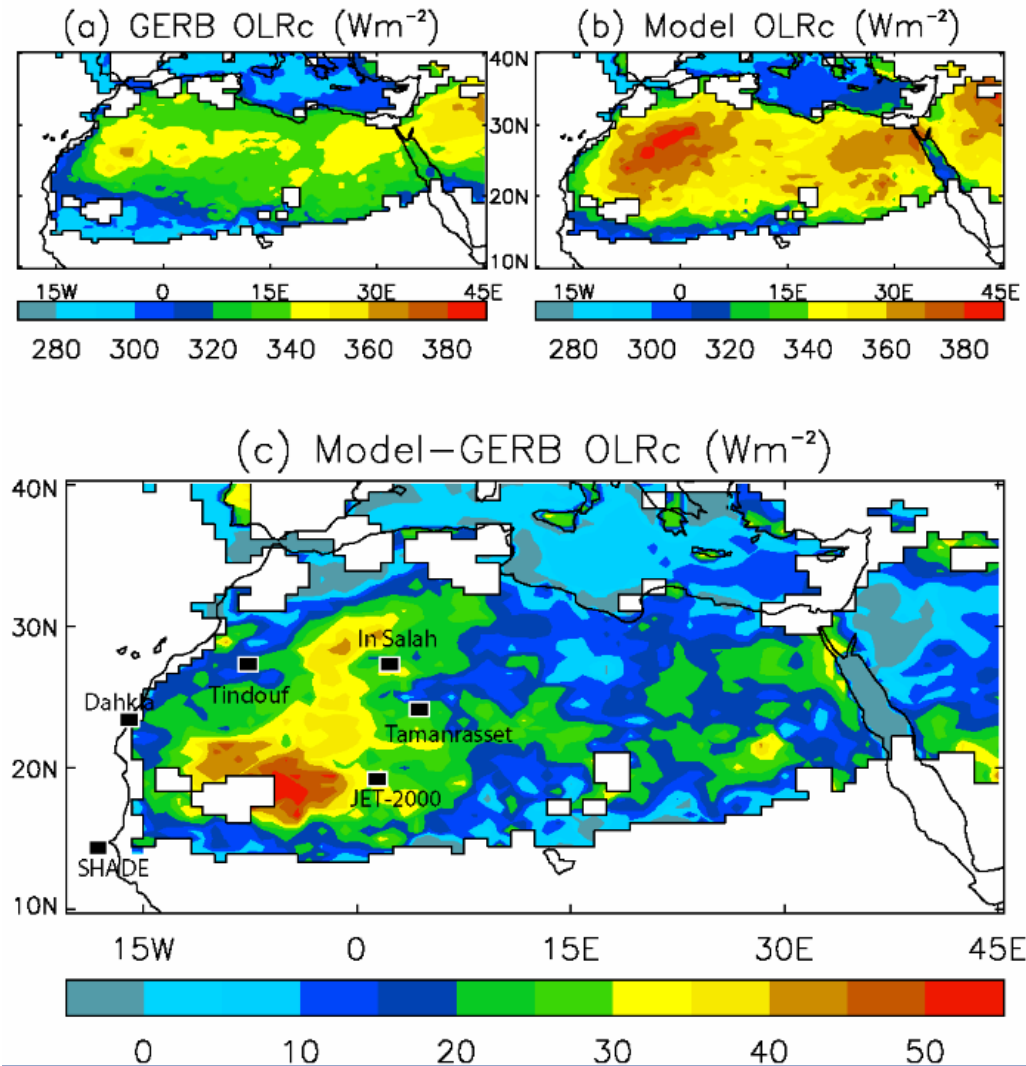
- They are part of the energy budget of the Earth: “aerosol-radiation interactions” or “direct radiative effect”.
- They affect visibility.
- For anthropogenic aerosols, they represent a radiative forcing of the climate system: RF_{aer} or direct radiative forcing.
- They allow model evaluation against remote-sensing products.

WHY WE CARE



- Forest fires in Portugal, 3 August 2003
- Image by MODIS radiometer
- Above dark surfaces, scattering by aerosols increase shortwave planetary albedo: loss of energy at the top of atmosphere and surface
- Above bright surfaces, absorption by aerosols decrease shortwave planetary albedo: gain of energy at the top of atmosphere, loss of energy at surface.

WHY WE CARE

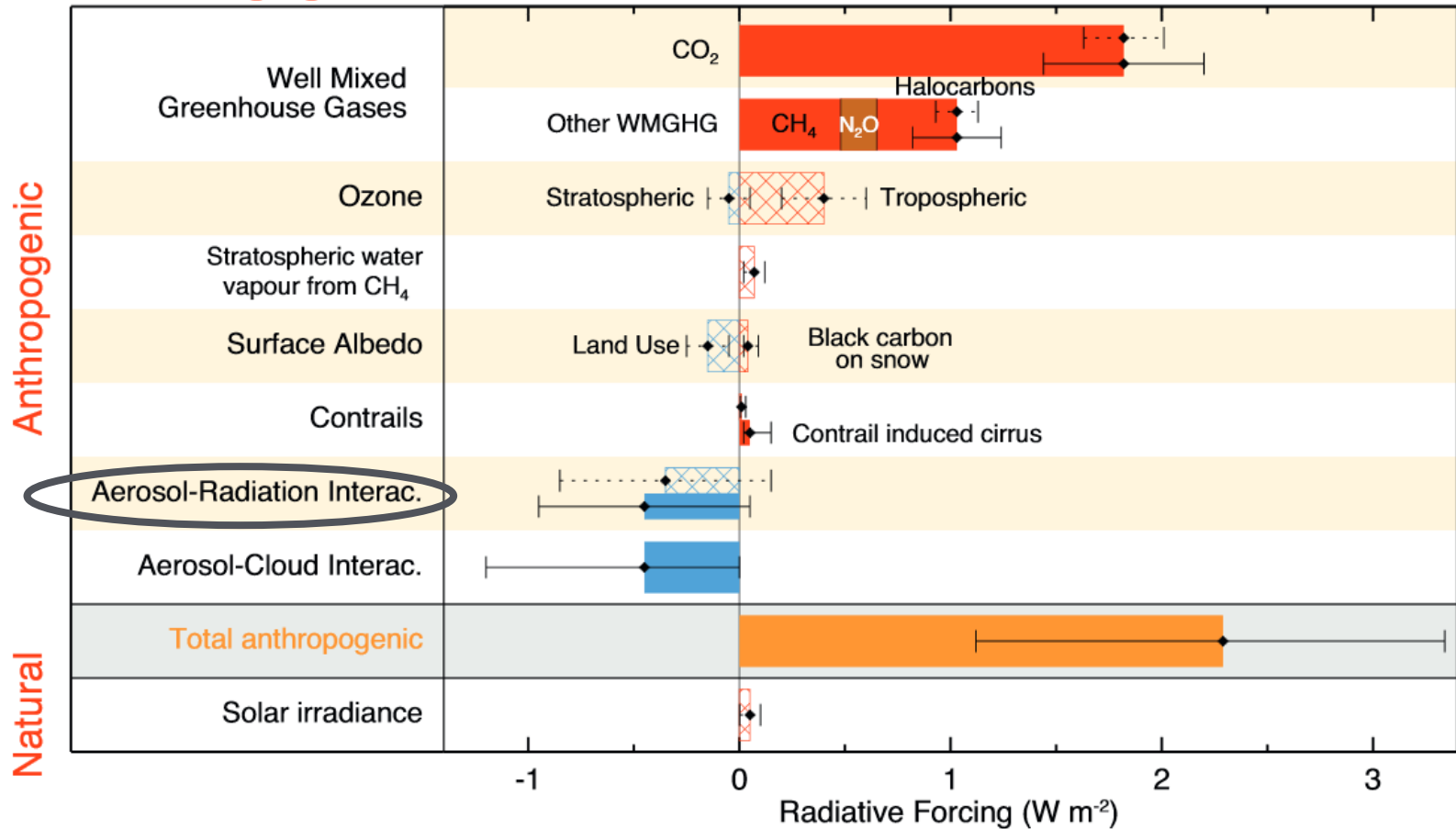


- Haywood et al. (2005) *Can desert dust explain the outgoing longwave radiation anomaly over the Sahara during July 2003?*

WHY WE CARE

Radiative forcing of climate between 1750 and 2011

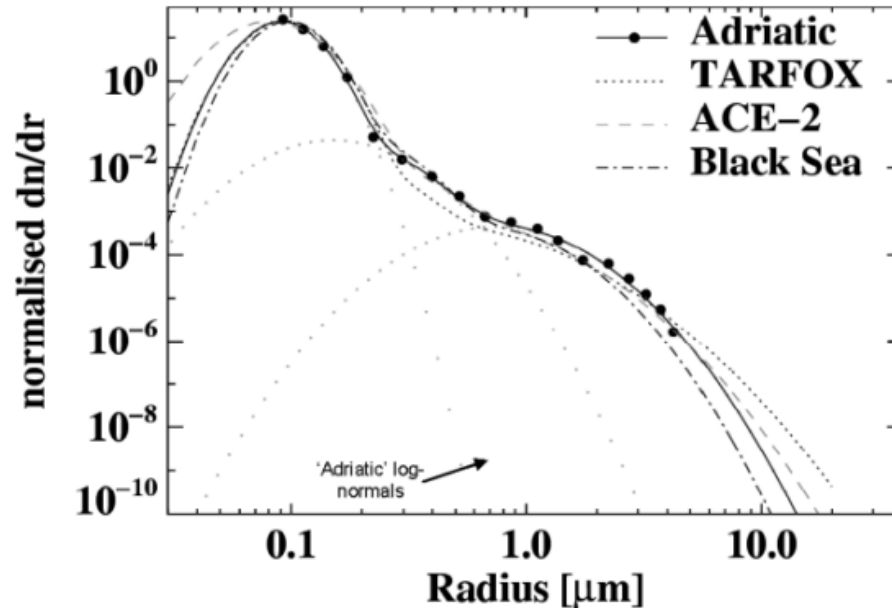
Forcing agent



IPCC AR5, Figure 8.15 (2013)

MIE SCATTERING

AEROSOL SIZE DISTRIBUTION



Aircraft measurements, Osborne *et al.* [2007]

- The distribution of particle number (or surface, or volume) as a function of particle radius shows local maxima, called modes.
- The size distribution is critical for interactions with radiation.

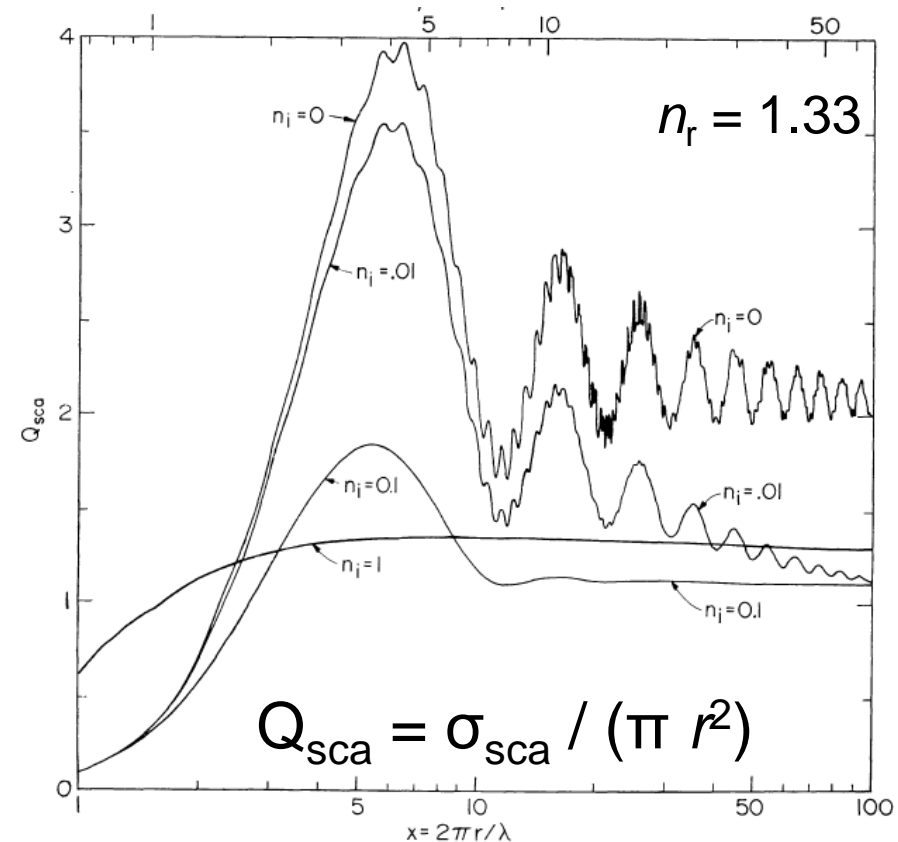
Radius range	$r < 0.05 \mu\text{m}$	$r < 0.5 \mu\text{m}$	$r > 0.5 \mu\text{m}$
Mode	Nucleation, Aitken	Accumulation	Coarse
Typical origin	Gas-to-particle conversion	Coagulation, combustion	Friction

MIE SCATTERING

- Aerosol radii r (0.1 to 10 μm) are of similar magnitude to the wavelength λ of shortwave and longwave radiation
 - Shortwave (or solar) spectrum: 0.25 to 5 μm
 - Longwave (or terrestrial) spectrum: 3 to 50 μm
- When $r \sim \lambda$, this is the domain of **Mie** scattering.
- Mie theory (1908) applies to homogeneous spheres, which is generally a good approximation for aerosols.
 - With the notable exception of mineral dust aerosols.

MIE SCATTERING

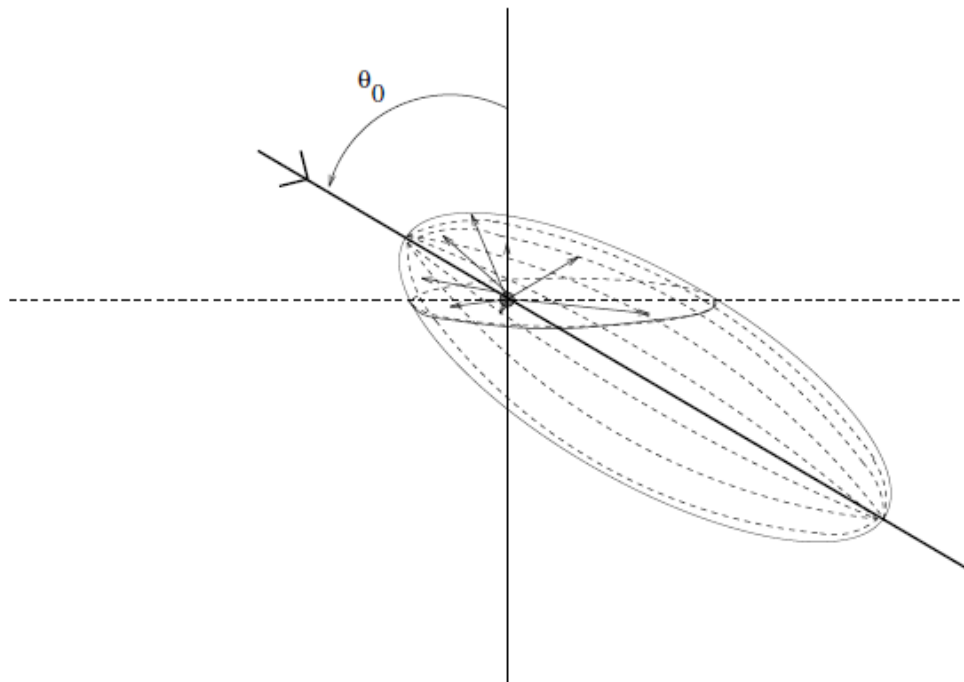
- Mie theory allows the calculation of the scattering and absorption cross sections σ_{sca} and σ_{abs} (in m^2).
- According to Mie theory, σ_{sca} and σ_{abs} depend *only* on:
 - The size parameter
 - $x = 2\pi r / \lambda$
 - The complex refractive index
 - $m = n_r - n_i$
- For hygroscopic aerosols, the impact of water uptake on x and m needs to be included.



Hansen and Travis (1974), Figure 9

PHASE FUNCTION

- Also calculated by Mie theory, the phase function gives the probability of being scattered in a given direction.



Boucher (2012)
Aérosols Atmosphériques
Figure 7.12

- In most climate models, phase function is represented by its average, the dimensionless asymmetry parameter g .

MIE SCATTERING FOR AN AEROSOL DISTRIBUTION

- The quantities valid for a given particle radius need to be integrated over the size distribution $n(r)$

$$\sigma_{sca}(r, \lambda, n_r, n_i) = \int Q_{sca}(r, \lambda, n_r, n_i) n(r) r^2 dr$$

- This is done for the three optical properties required:
- σ_{sca} , σ_{abs} , and g

RADAER

RADAER

Aerosol mass
Aerosol number
Aerosol dry diameter
Aerosol wet diameter
Aerosol density
Aerosol volume fraction, including water

UKCA/GLOMAP

Mie look-up tables
Refractive indices

UKCA_RADAER

Aerosol optical properties,
averaged over
SW and LW wavebands

Aerosol optical depths

SOCRATES
radiation scheme

RADAER: LOOK-UP TABLES

- Mie scattering calculations are too expensive to be done at runtime.
- However GLOMAP size distributions are interactive, so aerosol optical properties cannot be prescribed offline.
- Solution: look-up tables containing optical properties for all realistic combinations of x and m .
- Aerosol size distributions are assumed lognormal, with fixed standard deviations depending on mode.



RADAER: LOOK-UP TABLES

- The look-up tables contain:
 - $\sigma_{\text{sca}}(x, n_r, n_i)$ and $\sigma_{\text{abs}}(x, n_r, n_i)$, in m^{-1} (normalised per unit volume)
 - g
 - aerosol volume fraction $V(x) = \frac{4}{3} \pi \int x^3 n(x) dx$ (norm. per unit vol.)
- 51 values of x , 51 values of n_r , 51 to 801 values of n_i
- Realistic ranges depend on the aerosol mode and the range of wavelength considered:

Mode	Spectrum	x	n_r	n_i
Accumulation	SW	$4 \cdot 10^{-3}$ to 32	1.25 to 2	0 to 0.6
	LW	$4 \cdot 10^{-6}$ to 2	0.50 to 3	10^{-9} to 1
Coarse	SW	0.3 to 48	1.25 to 2	0 to 0.6
	LW	$3 \cdot 10^{-4}$ to 3	0.5 to 3	10^{-9} to 1

RADAER: LOOK-UP TABLES

- RADAER is set in the run_ukca namelist
- Rose: um > namelist > UM Science Settings > Section 34 UKCA

 l_ukca_radaer 	<input checked="" type="checkbox"/> true
Direct effect of MODE aerosols in radiation scheme	
 l_ukca_aie1 	<input checked="" type="checkbox"/> true
1st Indirect Effect of MODE aerosols (on radiation)	
 ukcaprec 	'\$UMDIR/vn\$VN/ctldata/spectral/ga7/RADAER_pcalc.ukca'
File of pre-computed values	
 l_ukca_aie2 	<input checked="" type="checkbox"/> true
2nd Indirect Effect of MODE aerosols (on precip.)	
 ukcaaclw 	'\$UMDIR/vn\$VN/ctldata/UKCA/radaer/nml_ac_lw'
LW file: aitken and insol acc modes	
 ukcaacsw 	'\$UMDIR/vn\$VN/ctldata/UKCA/radaer/nml_ac_sw'
SW file: aitken and insol acc mode	
 ukcaanlw 	'\$UMDIR/vn\$VN/ctldata/UKCA/radaer/nml_an_lw'
LW file: soluble accumulation mode	
 ukcaansw 	'\$UMDIR/vn\$VN/ctldata/UKCA/radaer/nml_an_sw'
SW file: soluble accumulation mode	
 ukcacrlw 	'\$UMDIR/vn\$VN/ctldata/UKCA/radaer/nml_cr_lw'
LW file: coarse-mode	
 ukcacrsw 	'\$UMDIR/vn\$VN/ctldata/UKCA/radaer/nml_cr_sw'
SW file: coarse-mode	
 l_ukca_radaer_sustrat 	<input checked="" type="checkbox"/> true
Sulphuric acid aerosol in stratosphere	

Here, vn10.4.

RADAER METHODS

1. Compute x from GLOMAP's modal wet diameter and current wavelength.
2. Compute m of internal mixture as volume-weighted average of component m (including water)
 - $[\text{NH}_4]\text{SO}_4$ or H_2SO_4 , BC, OC, SOA, sea-salt, mineral dust, NH_4NO_3
3. Now that x and m are known, access the right σ_{sca} , σ_{abs} , g , and V in the relevant look-up table.
4. Convert to specific scattering and absorption ($\text{m}^2 \text{kg}^{-1}$) to comply with radiation code requirements: $k_{\text{sca}} = \frac{\sigma_{\text{sca}}}{\rho V}$, where ρ is the modal density (kg m^{-3}).

RADAER

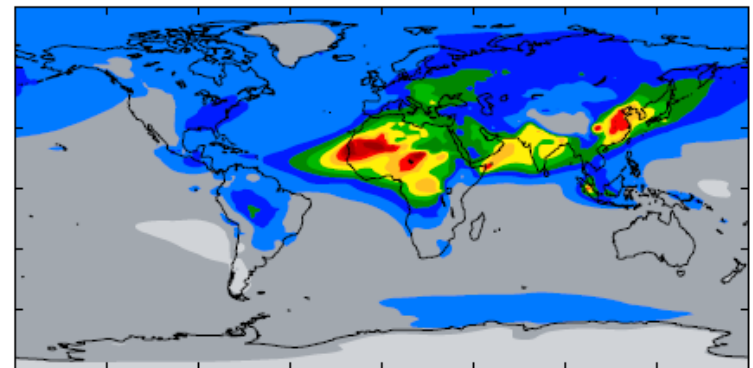
KEY FUNCTIONS

- `ukca_radaer_band_average()`
 - Integrates aerosol optical properties across Unified Model spectral wavebands
 - 6 integration wavelengths per waveband
 - Weighted by solar irradiance (SW) or Planck's blackbody function (LW)
- `ukca_radaer_compute_aod()`
 - Optical properties remain monochromatic (6 wavelengths)
 - Integration over the column

RADAER: DIAGNOSTICS

- Aerosol optical depths (AOD)
 - 1 per aerosol mode
 - Section 2 diagnostics (2–300+), with 6 pseudo-levels representing wavelength: 0.38, 0.44, 0.55, 0.67, 0.87, and 1.02 μm
- Allow comparison against satellite retrievals (MODIS, MISR, POLDER, ...) and ground-based measurements (AERONET)
- Angstrom exponent can be computed from AODs at two different wavelengths.

GLOMAP-mode AOD



Mean: 0.118



0 0.1 0.2 0.3 0.4 0.5

0.55 μm , vn7.3, Bellouin *et al.*, 2013.

RADAER: DIAGNOSTICS

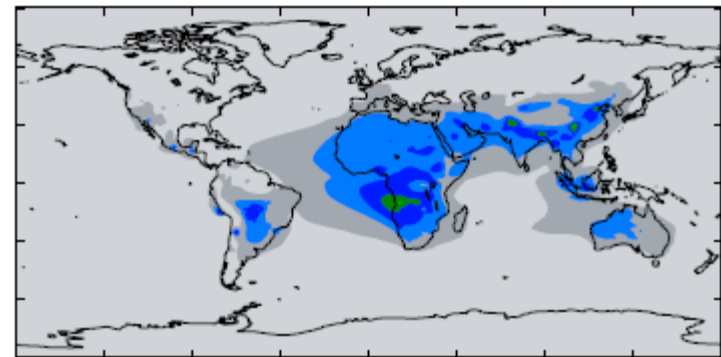
- Aerosol absorption optical depth (AAOD)
 - 1 per aerosol mode
 - Section 2 diagnostics, *require a branch*.
 - Same 6 wavelengths (pseudo-levels) as AOD

- Allow calculation of single-scattering albedo (SSA):

$$\varpi_0 = \frac{1 - AAOD}{AOD}$$

- Characterises absorption.

GLOMAP-mode SSA



Mean: 0.98



0.8 0.84 0.88 0.92 0.96 1

0.55 μm, vn7.3, Bellouin *et al.*, 2013.

RADAER: DIAGNOSTICS

- Other diagnostics available with branches:
 - Stratospheric aerosol optical depths
 - Same as AODs, but computed over stratospheric levels only.
 - Useful for volcanic eruption or geo-engineering studies.
- Vertical profile of scattering or absorption coefficients
 - Essentially, 3D profiles of AOD at fixed wavelengths (0.55 or 1.02 μm).
 - When divided by geometric thickness of model levels, give σ_{sca} or σ_{abs} , in m^{-1} .
 - Useful to compare against lidars.

SUMMARY

- RADAER is a piece of code within the Unified Model's SOCRATES Radiation Scheme.
- It allows interaction between radiation and UKCA/GLOMAP aerosols.
- It offers important diagnostics, e.g. aerosol optical depths.
- More details: Unified Model Documentation Paper 84, section 13.2
- Reference: Bellouin *et al.*, *Atmos. Chem. Phys.*, doi:10.5194/acp-13-3027-2013, 2013. [Section 5]
- n.bellouin@reading.ac.uk

UKCA_RADAER: DOUBLE-CALL SIMULATIONS

- Typically used to compute **radiative forcing** following the IPCC AR4 definition: “the change in net irradiance at the tropopause [...] with surface and tropospheric temperatures and state *held fixed at the unperturbed values.*”

